

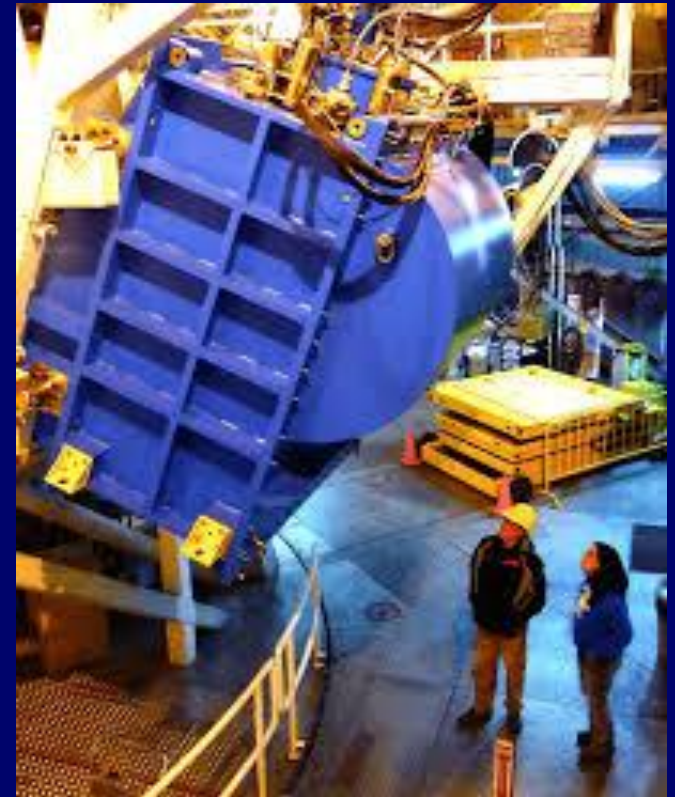
BISTRO – B-fields In Star-forming Region Observations



Pattle et al., 2015,
MNRAS, 450, 1094

Professor Derek Ward-Thompson
Director, Jeremiah Horrocks Institute
University of Central Lancashire
Cold Cores 2016, Budapest June 2nd

A SCUBA2-POL2 Survey on JCMT



The BISTRO Team

6 Co-PI's:

D Ward-Thompson (UK)

D Li (China)

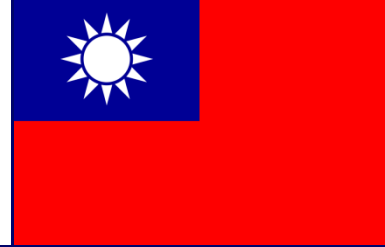
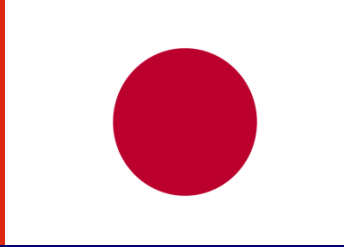
R Furuya (Japan)

R Rao (Taiwan)

W Kwon (Korea)

P Bastien (Canada)

+ ~100 co-I's from the above 6 countries.



Pierre Bastien
 Mike Chen
 Simon Coude
 James Di Francesco
 Jason Fiege
 Rachel Friesen
 Martin Houde
 Doug Johnstone
 Kevin Lacaille
 Brenda Matthews
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 Nagayoshi Ohashi
 Takashi Onaka
 Tae-Soo Pyo
 Hiroko Shinnaga
 Motohide Tamura
 Kohji Tomisaka

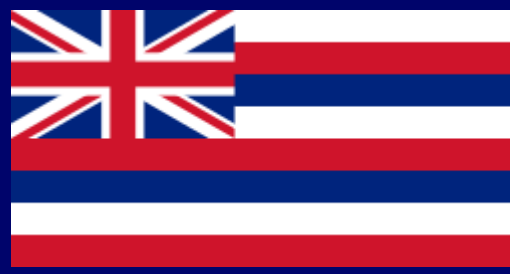
Do-Young Byun
 Jungyeon Cho
 Minhoo Choi
 Eun Jung Chung
 Il-Gyo Jeong
 Ji-hyun Kang
 Miju Kang
 Sung-ju Kang
 Gwanjeong Kim
 Jongsoo Kim
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 Jean-Francois Robitaille
 Giorgio Savini
 Anna Scaife
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David Berry
 Per Friberg
 Sarah Graves

Harriett Parsons
 Mark Rawlings



Sam Falle
 Sven van Loo
 Joe Mottram
 Sarah Sadavoy
 Yusuke Tsukamoto



BISTRO: Overview

- Aims to map Gould Belt star-forming regions in polarised light
- Awarded 224 hours of Band 2 observing time
- 100 survey members across 6 partner regions + EAO
- P.I.s: Derek Ward-Thompson (UK), Di Li (China), Ray Furuya (Japan), Woojin Kwon (Korea), Ramprasad Rao (Taiwan), Pierre Bastien (Canada)
- We aim to map the high-column-density regions of:

Ophiuchus, Orion A & B, Perseus, Serpens Main, Taurus L1495/B211, Auriga, IC5146

BISTRO: Scientific Goals

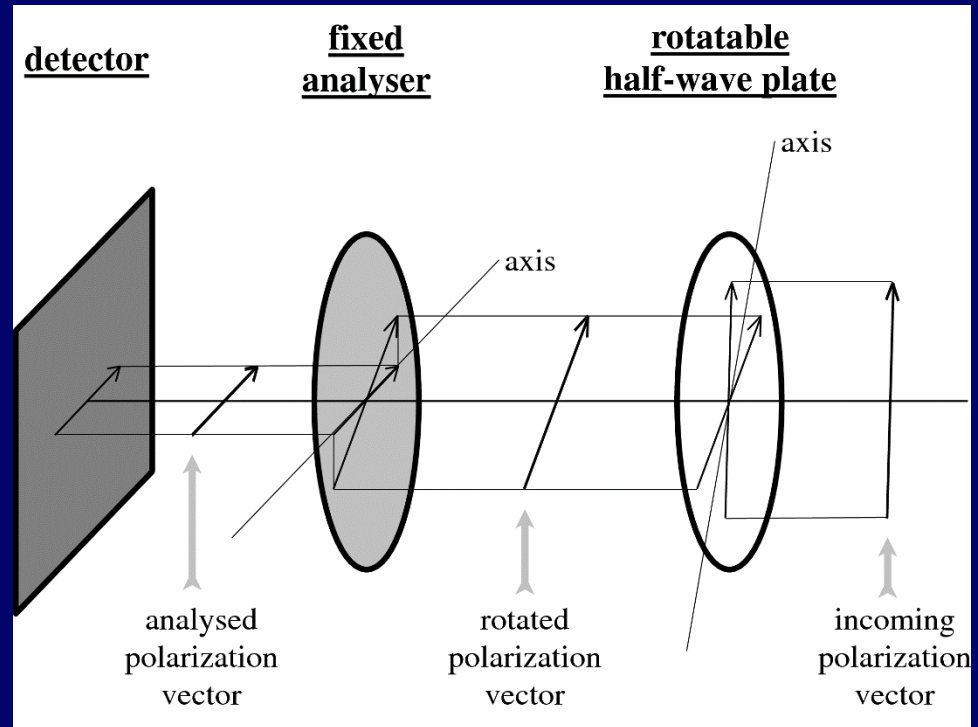
- To map the magnetic field within cores and filaments, on scales of ~ 1000 -5000 AU
- To determine magnetic field strengths in nearby molecular clouds
- To investigate the relative importance of magnetic fields and turbulence to star formation
- To test the model of magnetic funnelling of material onto filaments
- To investigate the role of magnetic fields in shaping protostellar evolution
- To investigate the effect of magnetic fields on bipolar outflows from young protostars

POL-2: The Instrument



A single-beam imaging polarimeter

Measures linear polarisation (Stokes Q & U)



Half-wave plate
(2Hz rotation)

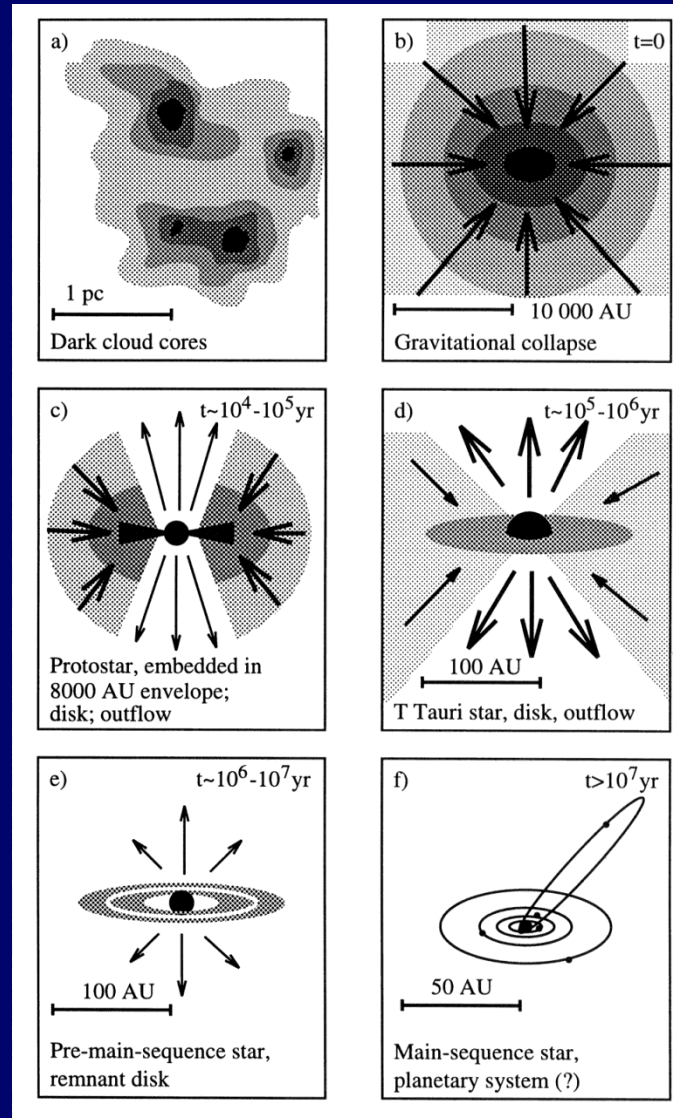
Fixed analysers

Credit: POL-2 User Manual

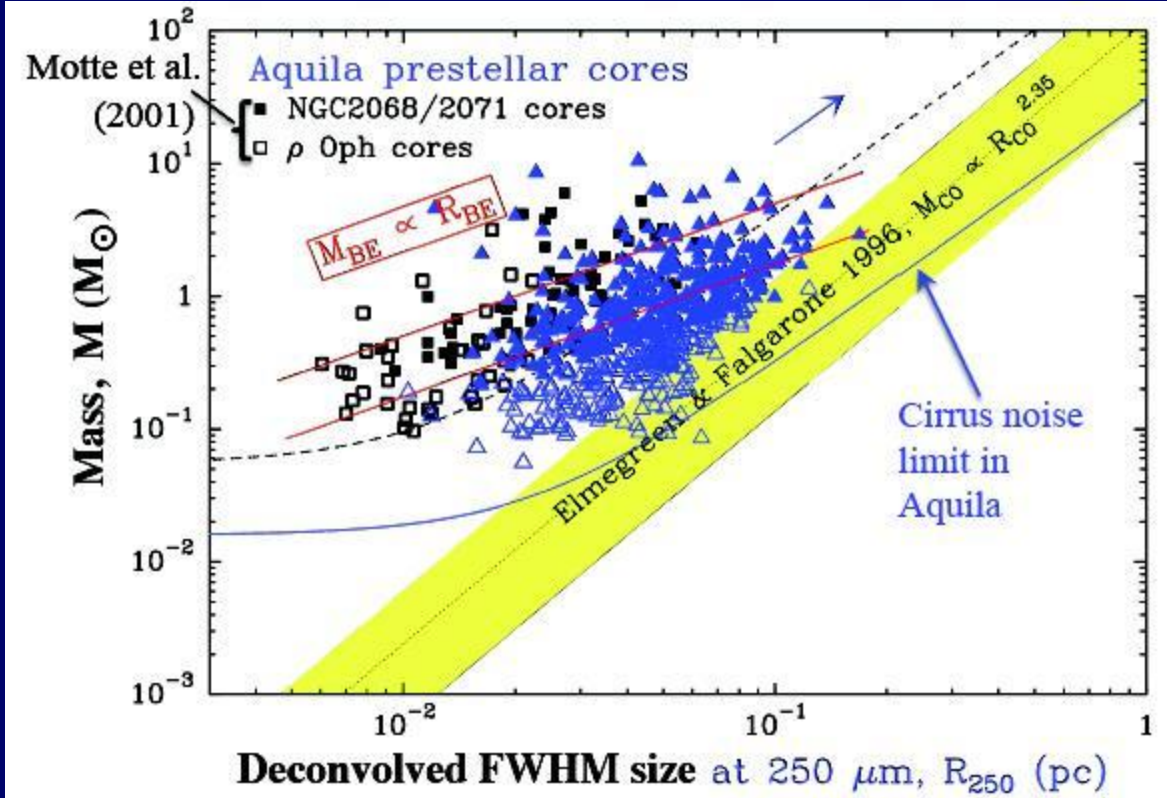
POL-2: Current Status

- Basic instrumental polarisation well-established:
1.3%, parallel to elevation axis, at 850 μ m
- Details of instrumental polarisation model being investigated:
Dependence on elevation
Variation across the focal plane
- Revised flux conversion factors: $\times 1.35$ at 850 μ m; $\times 1.96$ at 450 μ m
- Tiling of observations to map larger regions currently under investigation
- Observing has begun !!

The initial conditions of star formation

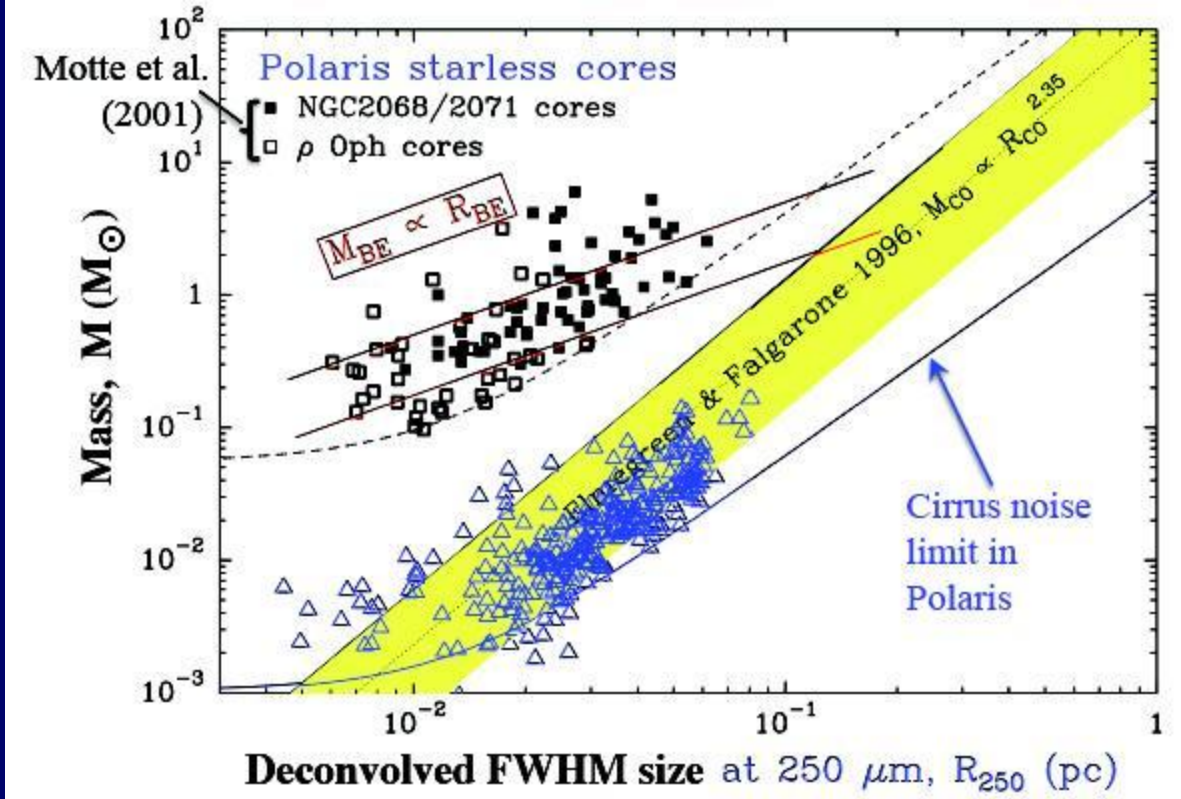


Are pre-stellar and starless cores the same population?



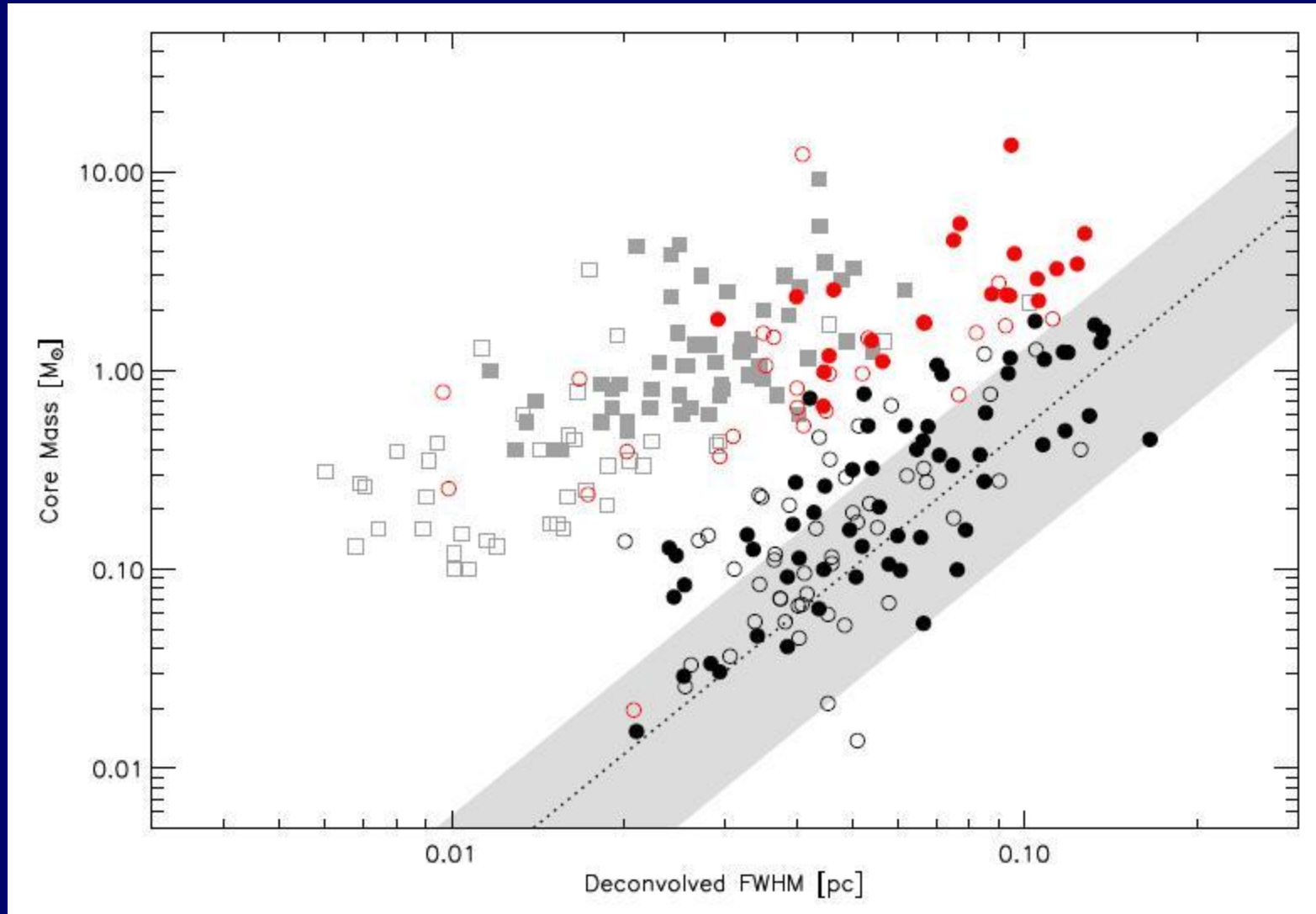
Andre et al., 2010, A&A, 518, L102

Are pre-stellar and starless cores the same population?



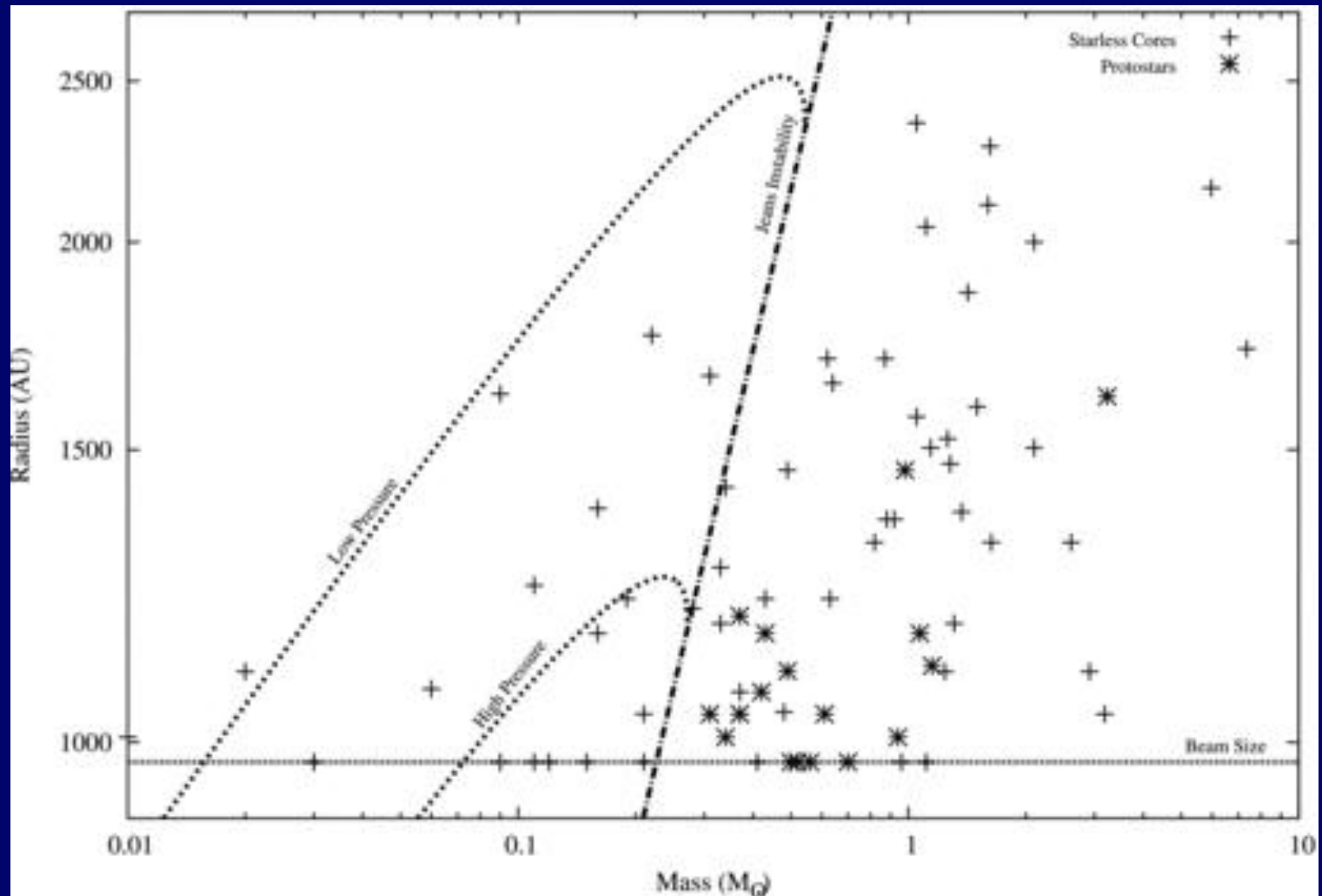
Andre et al., 2010, A&A, 518, L102

Seems to be just one population



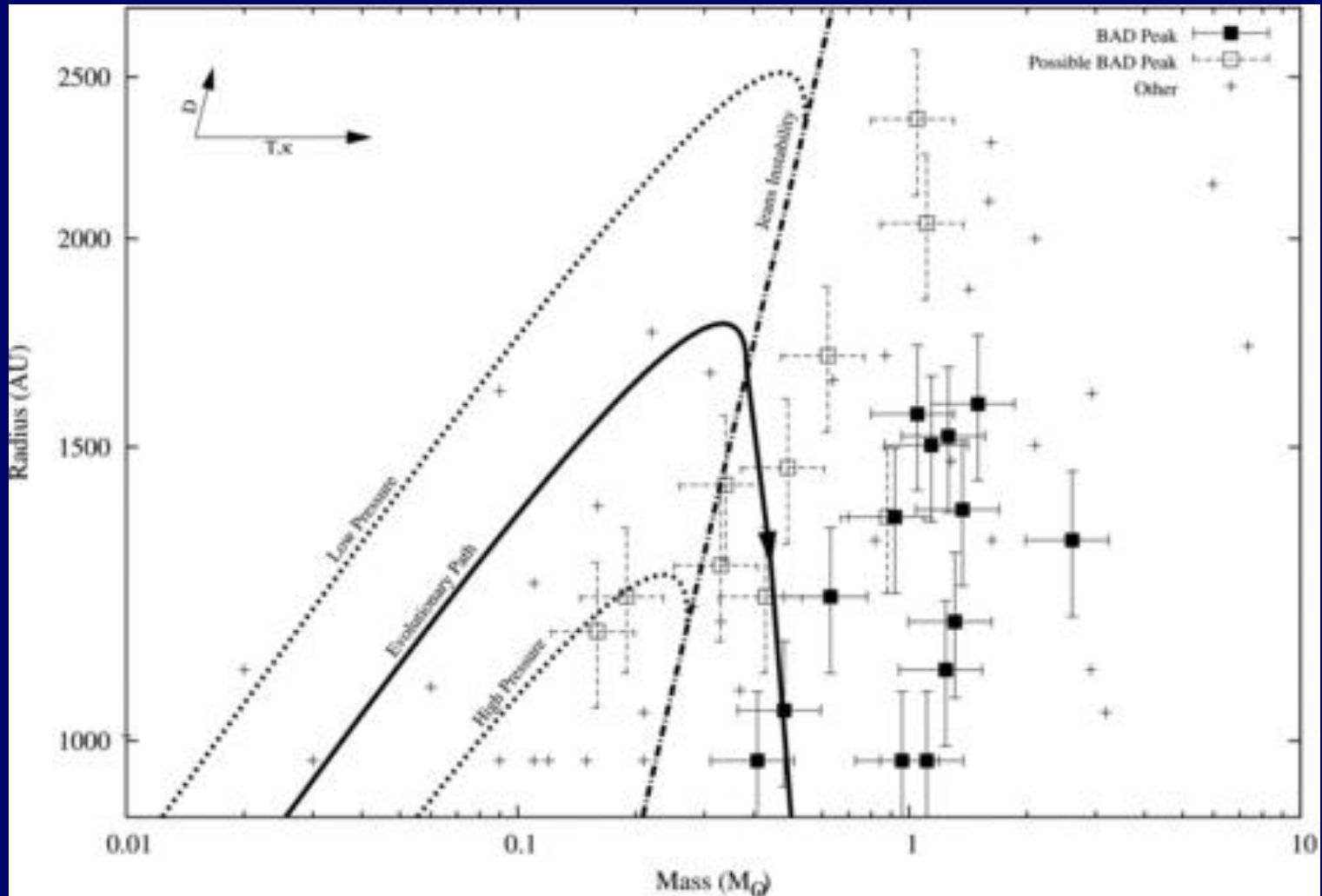
Kirk et al., 2013, MNRAS, 432, 1424

Starless and protostellar cores from Simpson et al.



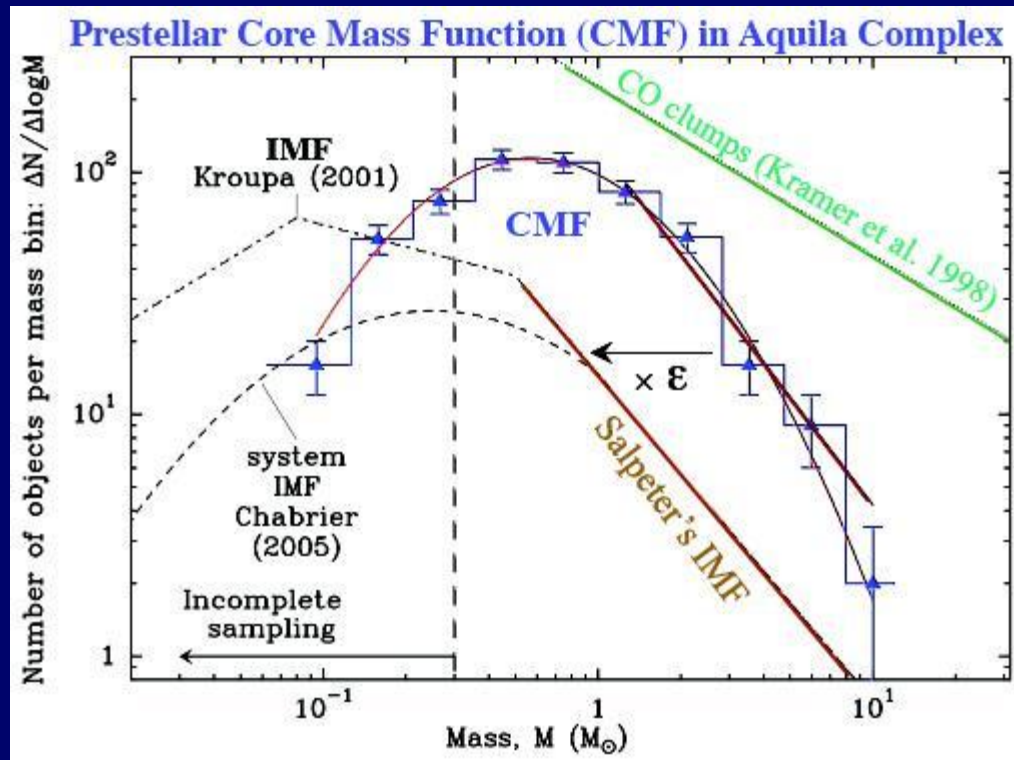
Simpson et al. MNRAS 2011;417:216-227

Same as Fig. 10, excluding the protostellar cores.



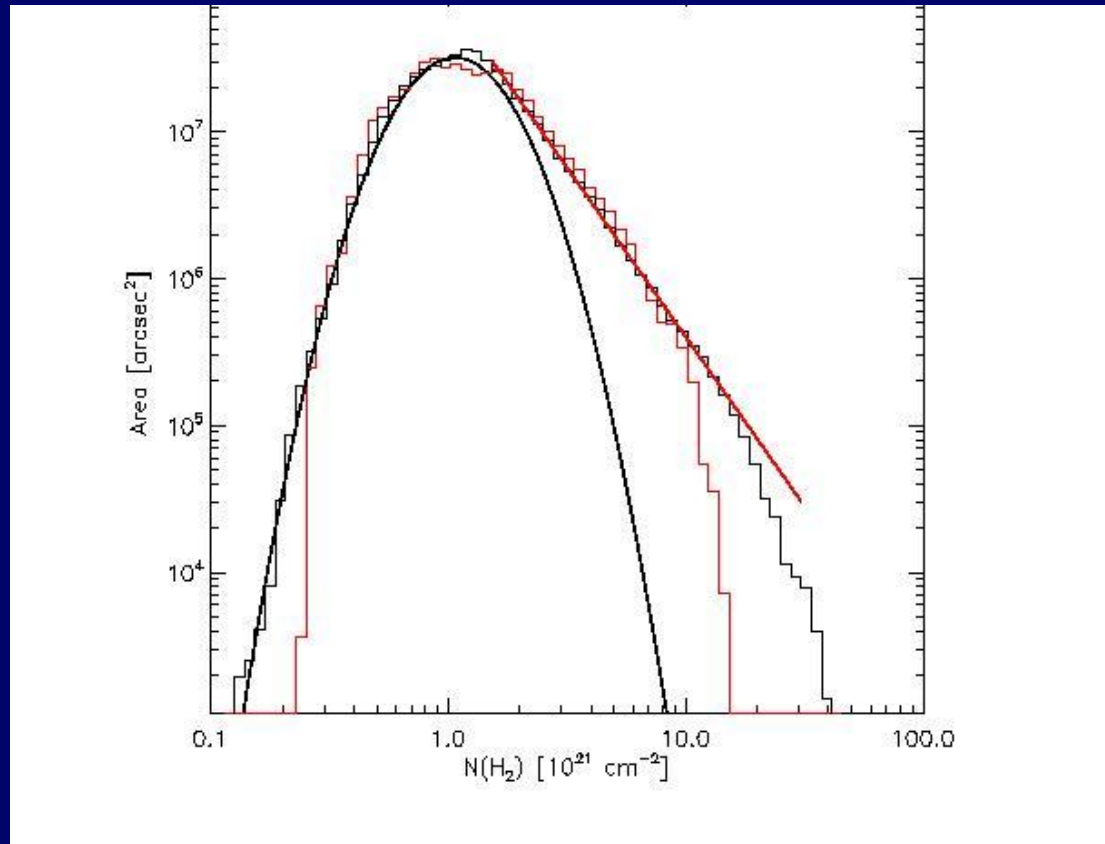
R. J. Simpson et al. MNRAS 2011;417:216-227

Prestellar/Starless CMF in Aquila



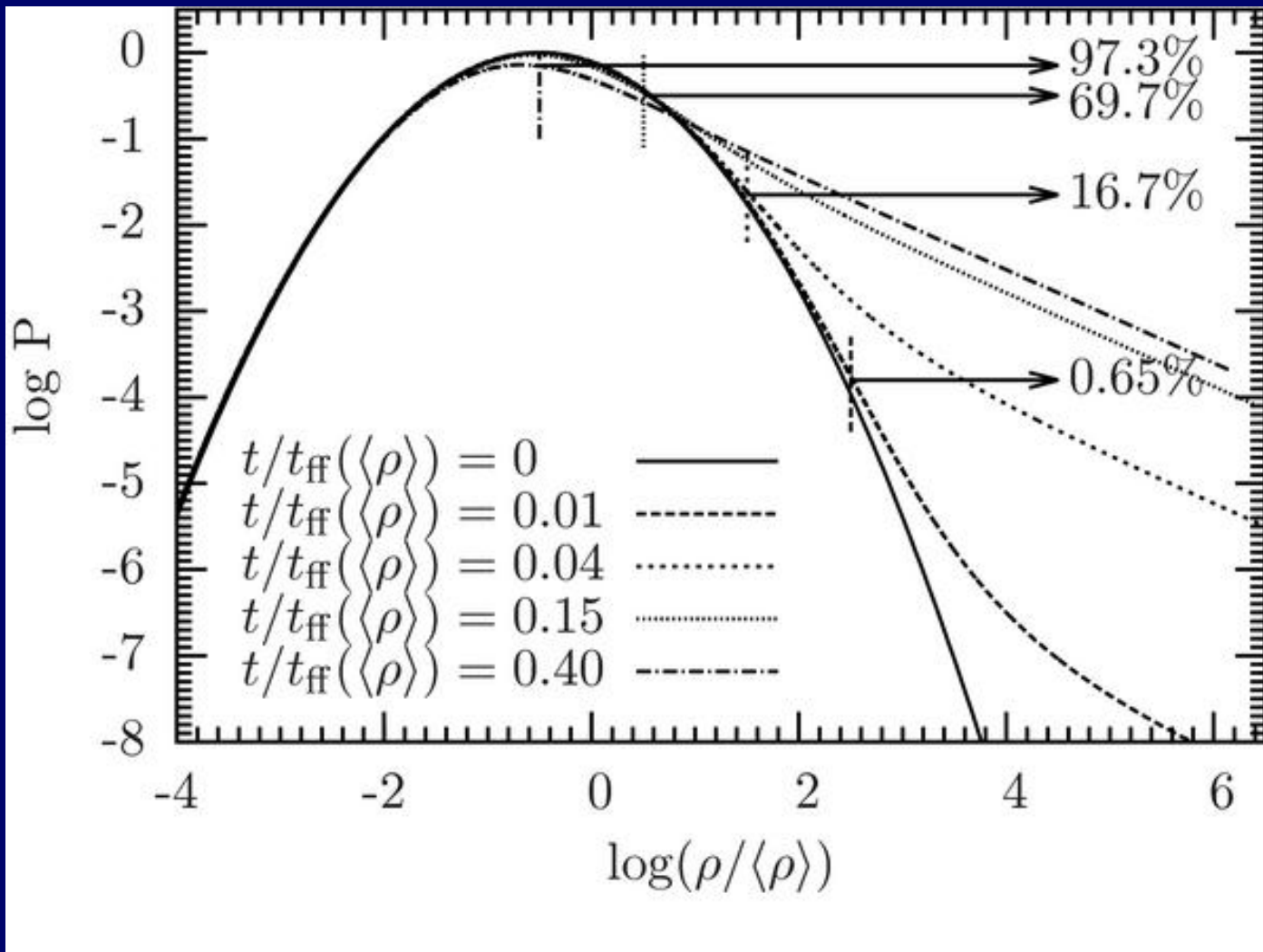
Andre et al., 2010, A&A, 518. L102

All of Taurus PDF



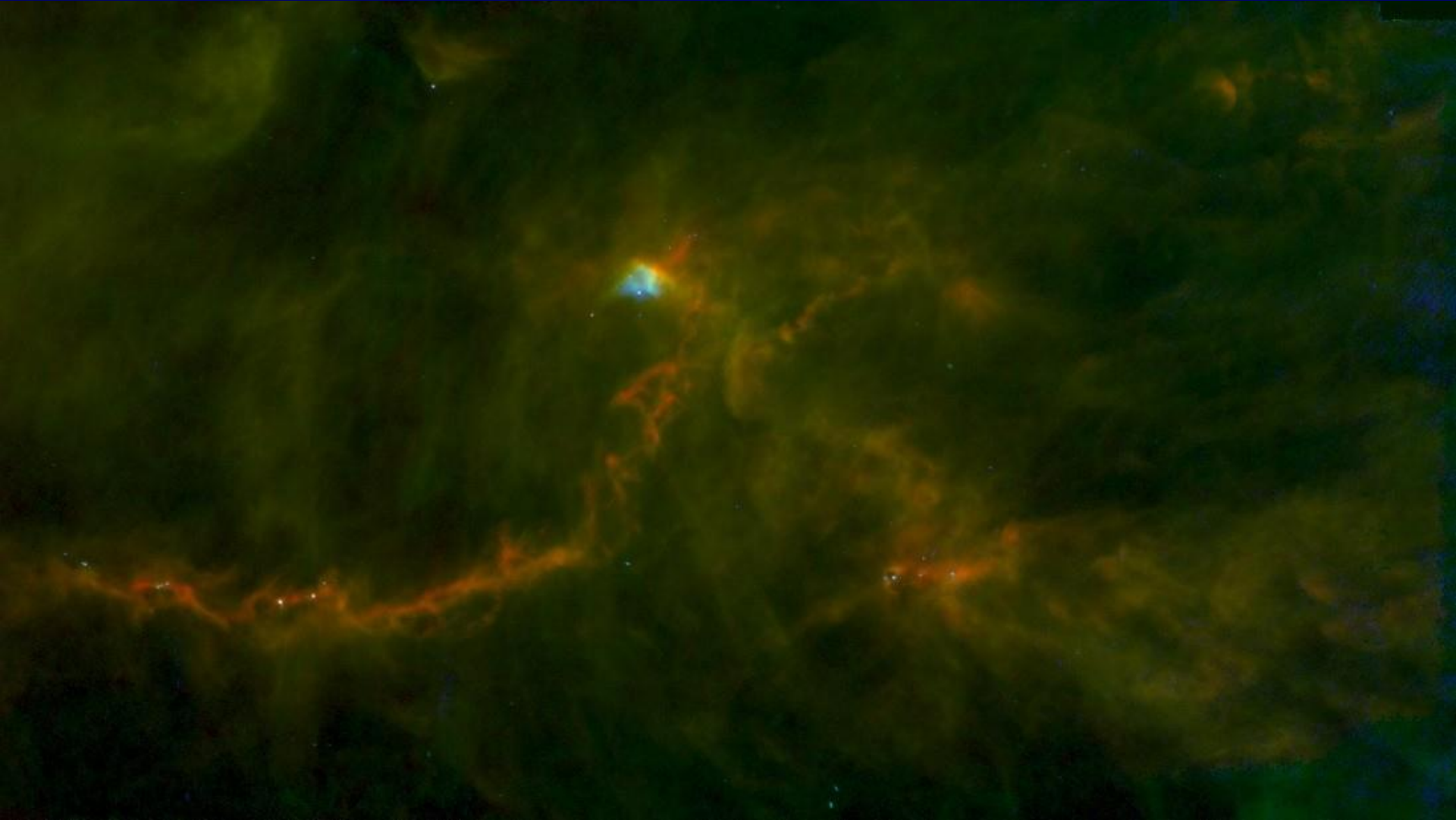
Kirk et al., 2016, in prep

Theoretical PDFs



Girichidis et al., 2014, ApJ, 781, 91

Taurus L1495 region: 70-160um



Marsh et al., 2014, MNRAS, 439, 3683
Marsh et al., 2016, MNRAS, 459, 342

L1495 central region

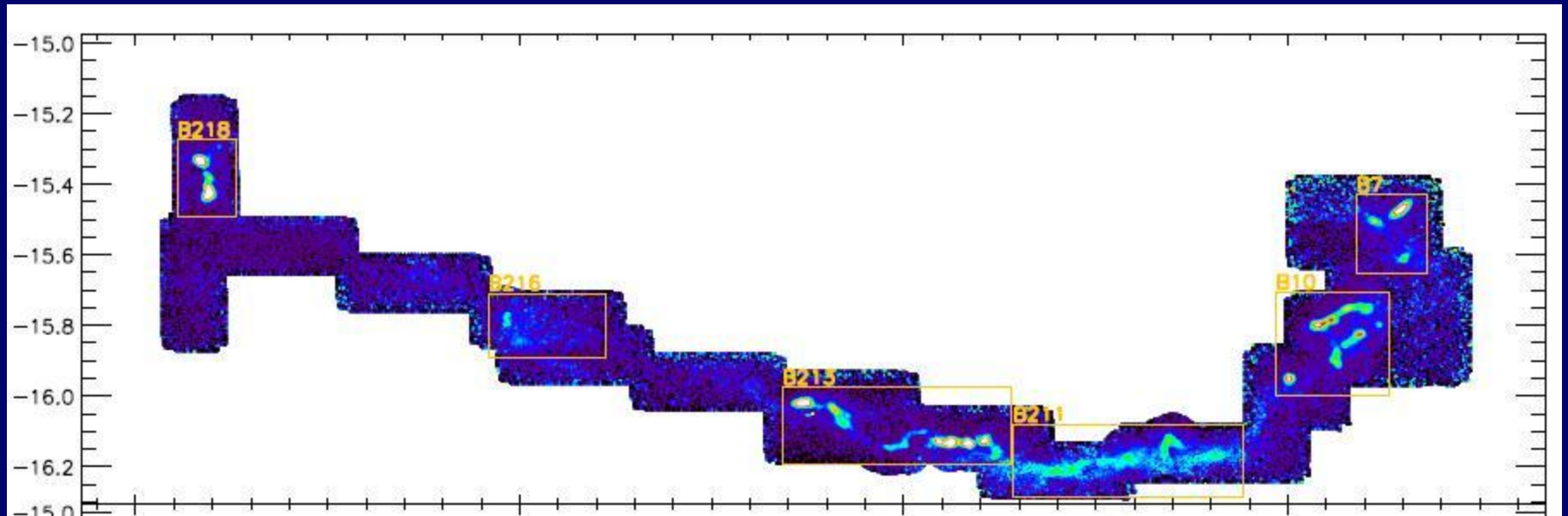
Red – 850um
Green – 500um
Blue – 250um



Ward-Thompson et al., 2016,
MNRAS, submitted

Note the red cores

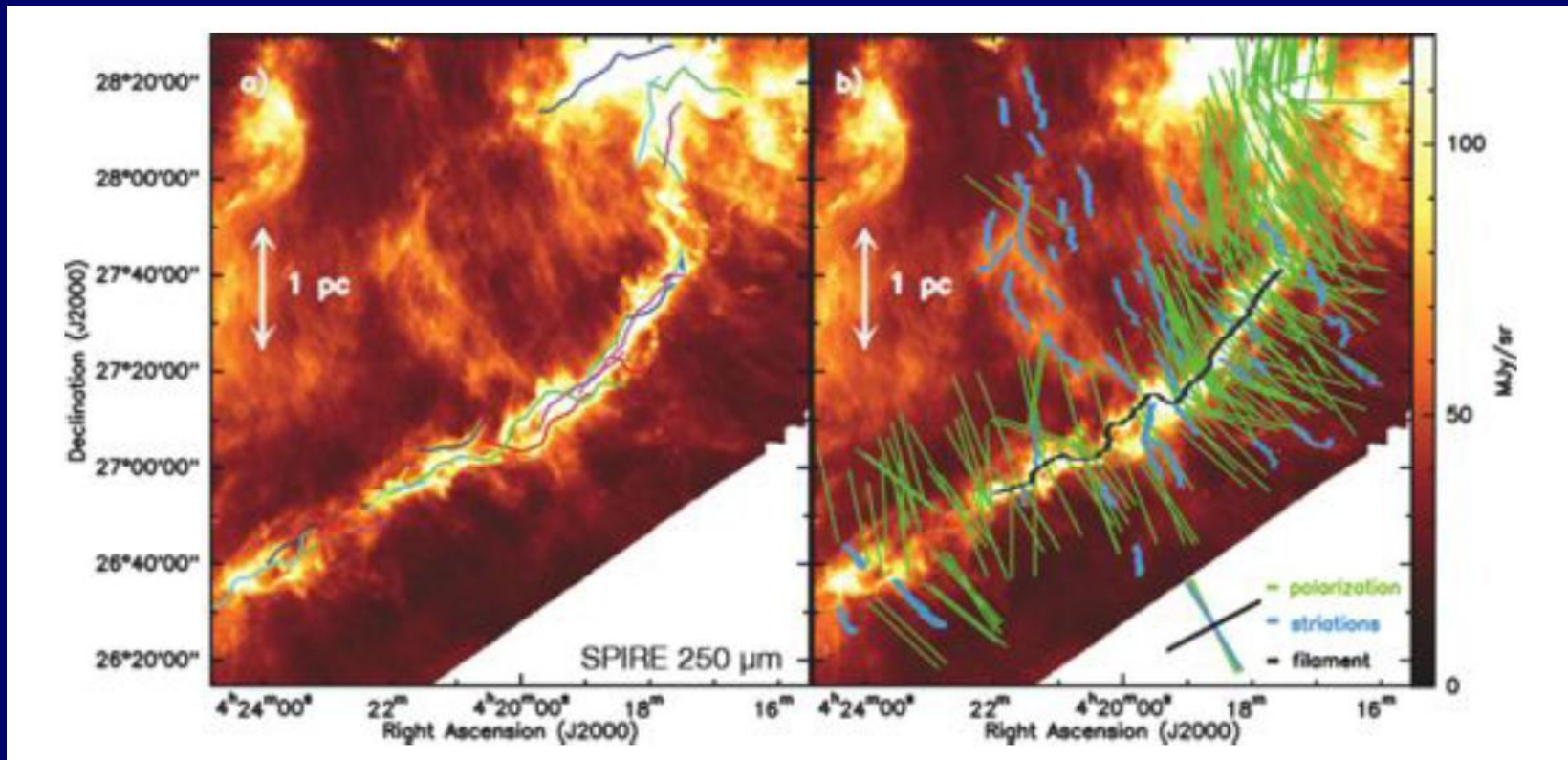
L1495 filament in NH₃(1,1)



Greenbank 100-m

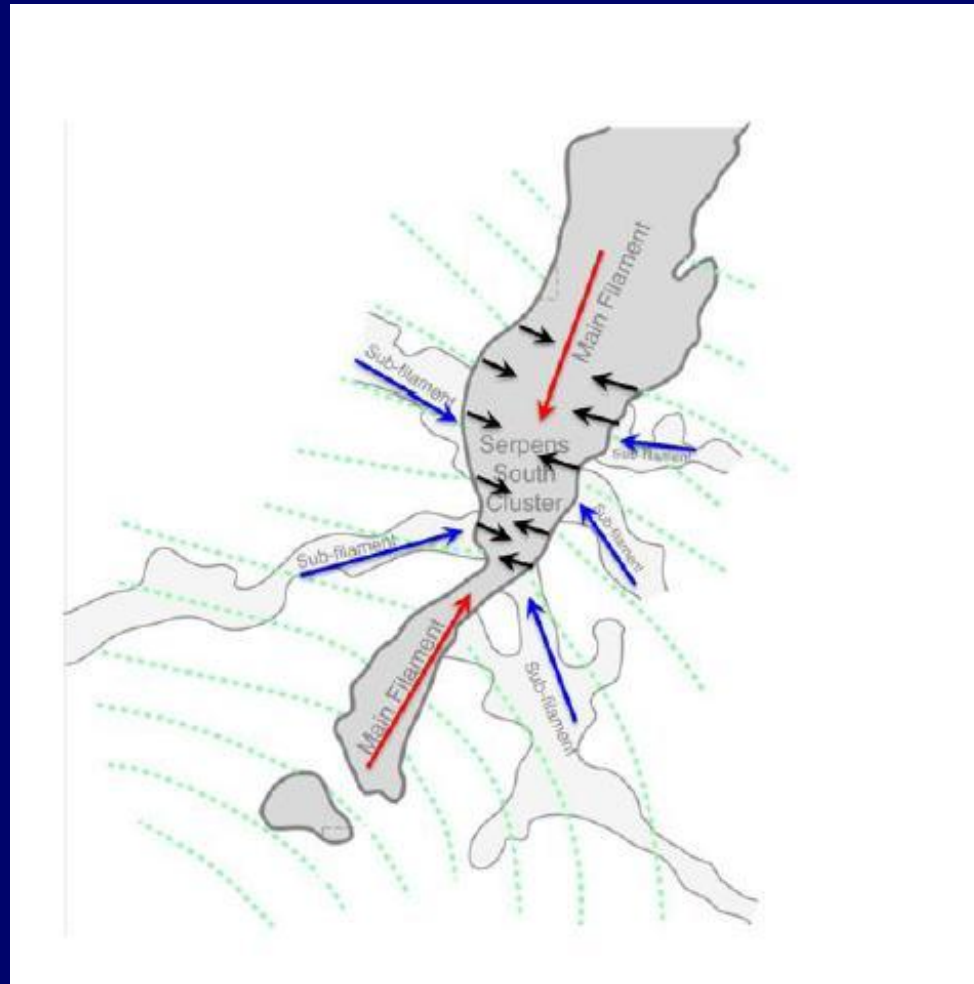
Seo et al., 2015, ApJ, 805, 185

Filament growth



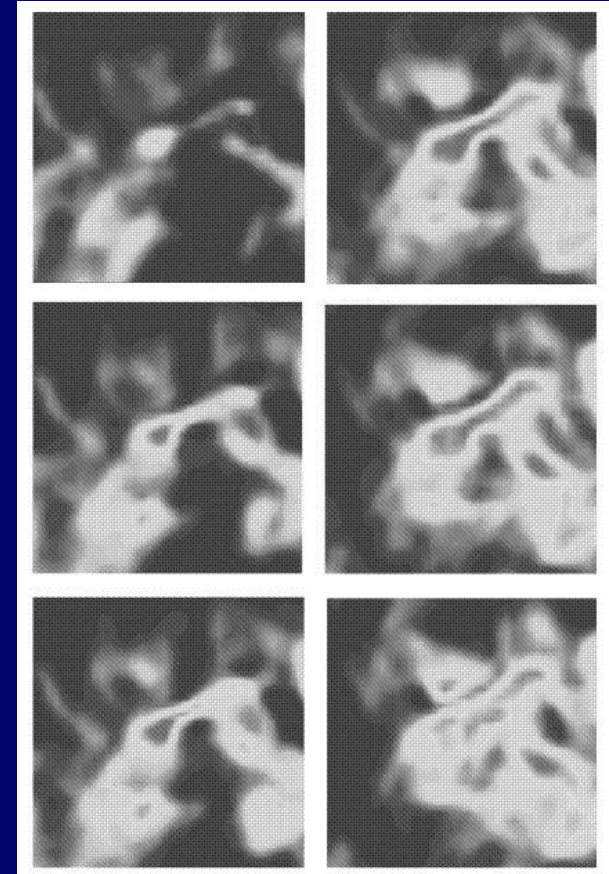
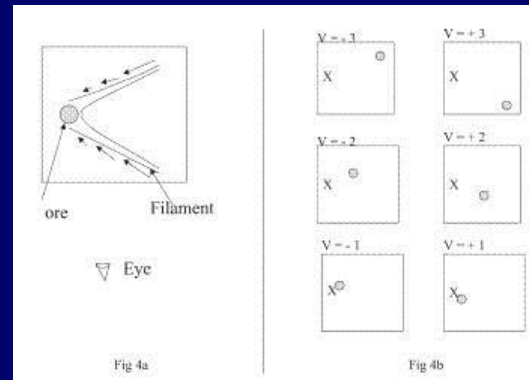
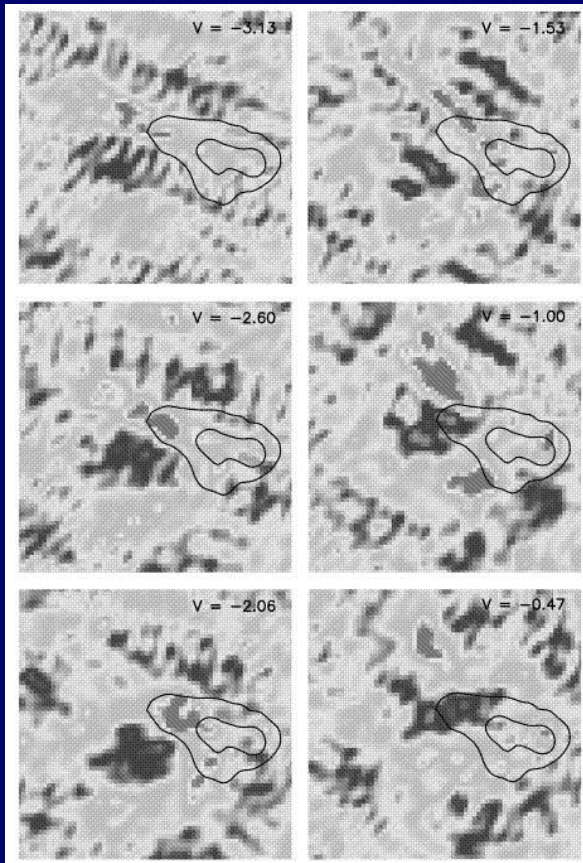
Palmeirim et al., 2013, A&A, 550, A38

Cores form on filaments

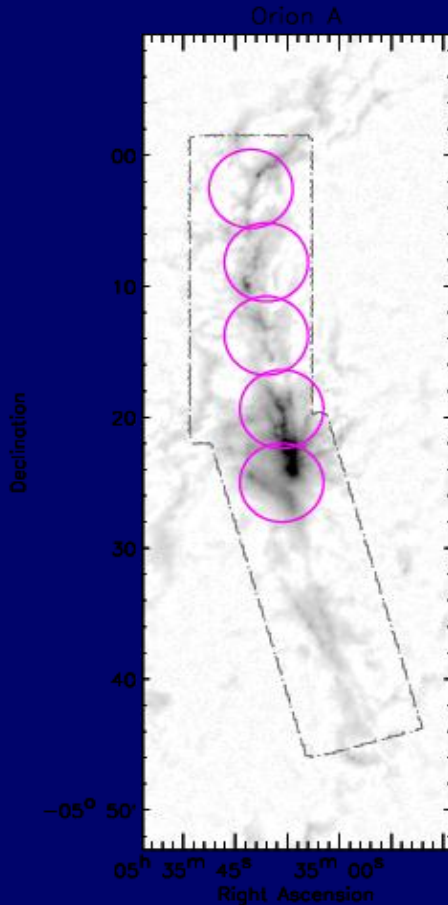


Andre et al., 2014,
PPVI, pp.27-51

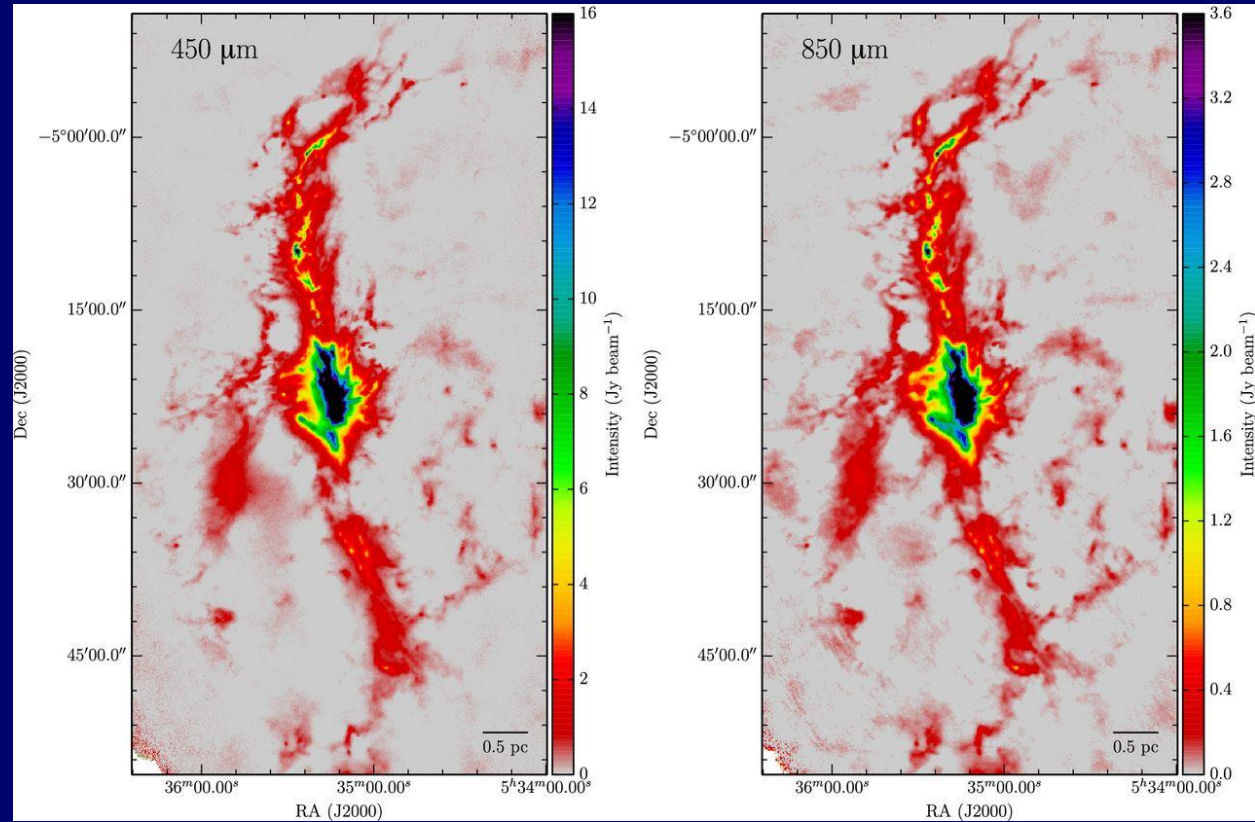
Filaments aren't new



Orion A Observations with JCMT



Credit: Kate Pattle



Salji et al., 2015, MNRAS, 449, 1769

THE REMAINDER OF
THIS TALK HAS BEEN
REDACTED

Conclusions

- The story so far:
 - Inter-stellar clouds rapidly become filamentary
 - Filaments form cores
 - Cores form stars
 - B-fields could be responsible for ‘funnelling’
 - B-fields are affected by outflows
- Future work:
 - What role do B-fields play inside filaments?
 - Are B-fields dominant?

SCUBA-2 GBS Papers

Ward-Thompson et al., 2007, PASP, 119, 855 – Survey description

Hatchell et al., 2013, MNRAS, 429, L10 – Perseus, NGC1333

Sadavoy et al., 2013, ApJ, 767, 126 – Perseus, B1

Dodds et al., 2015, MNRAS, 447, 722 – Perseus, NGC 1333, discs

Rumble et al., 2015, MNRAS, 448, 1551 – Serpens, MWC 297

Salji et al., 2015, MNRAS, 449, 1769 – Orion A North, cores

Salji et al., 2015, MNRAS, 449, 1782 – Orion A North, filaments

Buckle et al., 2015, MNRAS, 449, 2472 – Taurus, L1495

Pattle et al., 2015, MNRAS, 450, 1094 – Ophiuchus

Mairs et al., 2015, MNRAS, 454, 2557 – Data reduction comparison

Kirk et al., 2016a, ApJ, 817, 167 – Orion B, cores

Kirk et al., 2016b, ApJ, 821, 98 – Orion B, clustering

Coudé et al., 2016, MNRAS, 457, 2139 – Orion A North

Chen et al., 2016, ApJ, in press, arXiv:160506136 – Perseus, grain growth

Rumble et al., 2016, MNRAS, in press, arXiv:160504842 – Serpens, W40

Broekhoven-Fiene et al., 2016, ApJ, submitted – Auriga

Mairs et al., 2016, MNRAS, submitted – Orion A South

Ward-Thompson et al., 2016, MNRAS, submitted – Taurus, L1495, Herschel comparison

Bresnahan et al., 2016, MNRAS, in prep – Corona Australis

Pattle et al., 2016, MNRAS, in prep – Cepheus

Lane et al., 2016, MNRAS, in prep – Orion A, clustering

Mowat et al., 2016, MNRAS, in prep – Lupus I

Ciccone et al., ApJ., in prep. – IC5146